may find a certain physical association existing among them notwithstanding their dissimilarity of orbit. The most significant fact in connection with this subject is that certain sharplydefined points exhibit a numerous retinue of showers, while the spaces immediately adjacent are comparatively barren of such displays. The radiant points coincide at particular centres, to the marked exclusion of closely-bordering regions. observe with the utmost fulness and accuracy the evidence which the sky affords, leading to this remarkable conclusion, and to discover, if possible, the meaning of this singular persistency of showers, will be the important aim of future observations; it may clear away a difficulty from observers, and perhaps enlarge our views as to the visible character of meteor systems. uninterrupted appearance of shooting-stars in the nocturnal sky offers the ready means of attacking the problem anew and removing any doubts which may still exist as to the stationary, long-enduring aspect of many showers, which must, indeed, remain an indelible effect of all full and trustworthy observation.

Bristol: 1884, October.

Observations of Comets Pons-Brooks and Ross. By A. B. Biggs.

(Communicated by the Secretaries.)

(By triangular-bar Micrometer. Dark field.)

Date.	Mean Time. (Launceston.)	Diff. R.A. Diff. Decl. Comet from Star.		Name of Star.	Hour Angle. (Approx.)	
1884. Jan. 26	h m 9 52 O	m s + I 33.25	+ 36 58.6	7 Ceti	h m	
	956 o	+ 1 36	+37 9.4	,,		
29	8 52 o	- 2 51 5	-22 30	106 Lacaille	5 2 W.	
	8 57 o	- 2 49.7	-21 59	,,	5 8	
	9 4 10	- 2 47.5	-2I o.8	,,	5 15	
Feb. II	8 45 23	- o 19.3	+ 28 59	305 Lacaille	$5 9\frac{1}{2}$	
	8 55 21	- o 18.3	+ 29 36	. ,,	$5 \ 19\frac{1}{2}$	
23	10 10 0	+ 3 15.7	+ 28 56	$oldsymbol{\gamma}$ Phenicis	$654\frac{1}{2}$	
	10 16 45	+ 3 16.2	+29 19.6	,,	7 0	
Mar. 2	9 53 0	- 3 30	- 12 57	542 Lacaille	7 0	
11	8 17 30	+ 8 32.7	-17 41	588 ,,	5 35	
12	7 35 30	- IO 5·5	+ 9 27	693 ,,	4 53	
13	8 13 o	- 7 1 5 5	+ 39 57	,1	5 34 .	
	8 28 o	- 7 18.5	+40 43.5		$5 \cdot 53\frac{1}{2}$	

Date. Mean Time. (Launceston.)			Diff. R.A. Diff. Decl. Comet from Star.			Name of Star.	Hour Angle. (Approx.)			
Apr. 1834	3	h m 8 48	45	_	m s O 4	- 11 36"	989 Lacaille	h m 6 38		
		8 57	45		0 3.3	- 10 I	***	6 47		
	4	8 50	20	+	3 33	+10.15	"	6 29		
		8 59	20	+	3 45.5	+ 10 34	"	6 48		
Comet "Ross."										
Feb.	1	9 15	О	+	o 14 [.] 7	+ 19 43	9623 Lacaille	6 14		
		9 25	0	+	0 17	+ 18 24	"	6 24		

A very hazy object; nebulous; measures very difficult; no definite point. The only opportunity afforded for obtaining measures.

In all the above measures, different refraction and the comet's proper motion are not reckoned for.

Launceston, Tasmania: 1884.

Ephemeris for Finding the Positions of the Satellites of Uranus, 1885. By A. Marth.

The angle of position P of the minor axes, the major and minor semi-axes a and b of the apparent ellipses described by the satellites, the longitudes u—U of the satellites reckoned in their orbits from the points which are in superior conjunction with the planet's centre and the planeto-centric latitude of the Earth above the assumed plane of the orbits, are approximately the following:

tho i	.0110 11 111	ъ.							
			Ariel.				Umbrie		
Greenw. no	on P.	a_1	b_1	$u_1 - U$	Diff.	a_z	b_2	$u_2 - U$	Diff.
1885.	٥,	11	#	0	0	"		T 2 4: O T	•
Jan. 13	285.45	14.75	+ 4.59	114.40	1428:46	20.54	+ 5.97	134.01	868.75
23	.46	14.87	4.30	103.16		20.72	5.99	282.76	.72
Feb. 2	·47	14.99	4.28	91.60	·41	20.89	5.97	71.48	.70
12	.48	15.09	4.54	80.01	.38	21.03	5.91	220.18	·69
22	.20	15.18	4.18	68.39	.35	21.14	5.83	8.87	. 66
Mar. 4	.51	15.24	4'10	56 [.] 74	.33	21.53	5.71	157.23	·65
14	.53	15.27	4.00	45.07	.30	21.27	5.58	306.18	.63
24	285.54	15.58	+ 3.89	33.37	·27	21.28	+ 5 43	94.81	.61
Apr. 3	.55	15.26	3.78	21.64	.26	21.26	5.52	243.42	·61
13	•56	15.21	3.67	9.90	.24	21.19	2.11	32.03	60
23	·5 7	15.14	3.26	358.14	.23	21.10	4 [.] 96	180.63	.60
May 3	.57	15.02	3.46	346.37	.23	20.97	4.82	329.53	.60
13	.57	14.95	3.37	334.60	.22	20.82	4.70	117.83	·60
23	•58	14.83	3.31	322.82	1428 22	20.66	4 [.] 61	266 [.] 43	868-61
June 2	285.57	4.70	+ 3.56	311.04	-	20.48	+ 4 54	55 ^{.0} 4	